



Multispectral Imaging Spectroscopic Equipment


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Spectroscopic Equipment

Light source	Analyzer	Detector
Sun Lamps Lasers Synchrotron	Spectrometers	Diodes Photo multiplier tube CCDs Image intensifiers

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Line light sources

Energy

Atom or Ion

Excited state lifetime

$$I = I_0 e^{-t/\tau}$$

Natural line width

$$\Delta\nu_n = \frac{1}{2\pi\tau}$$

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Doppler broadening

Velocity in viewing direction

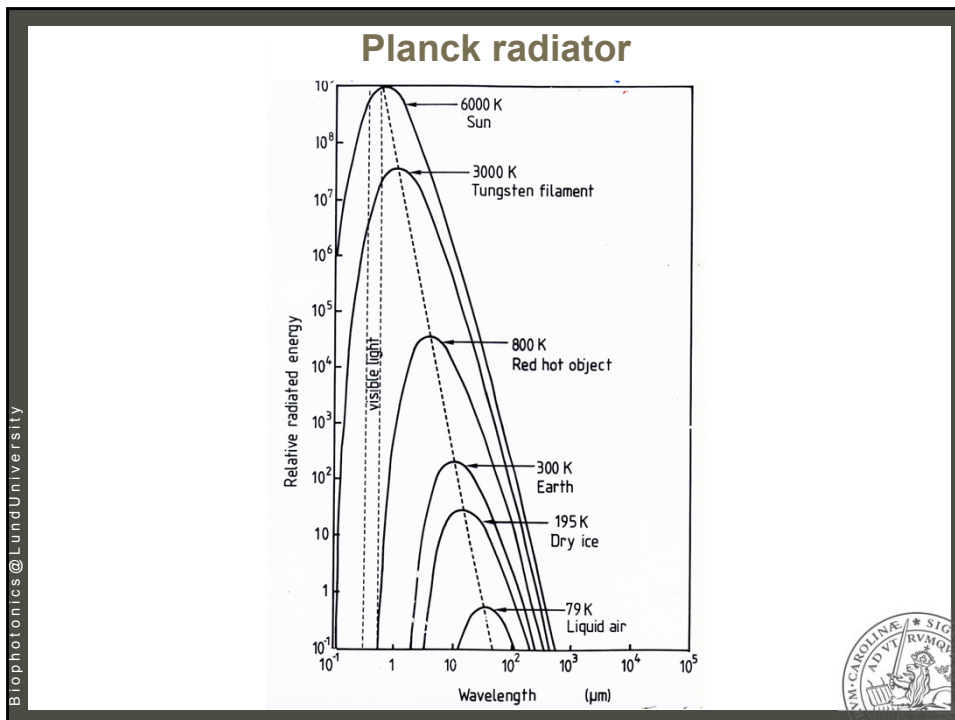
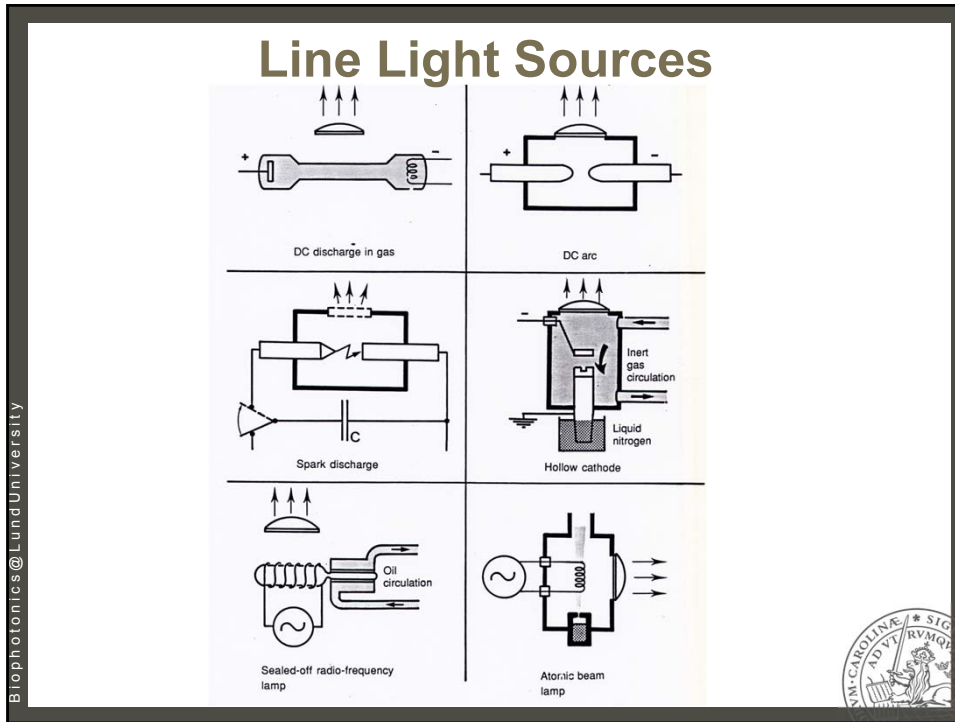
Doppler width

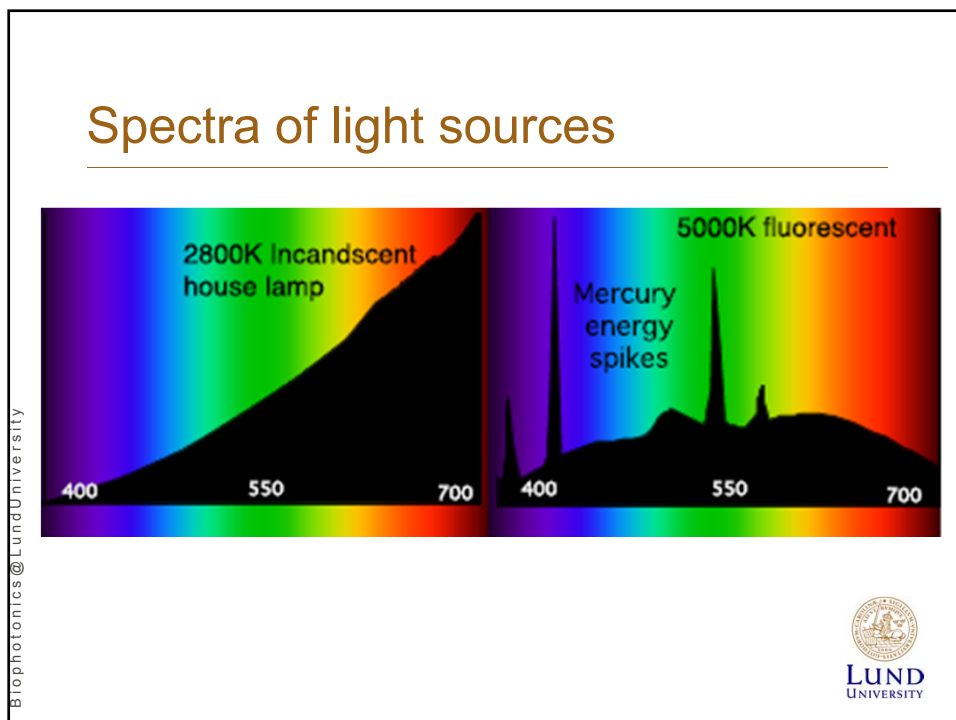
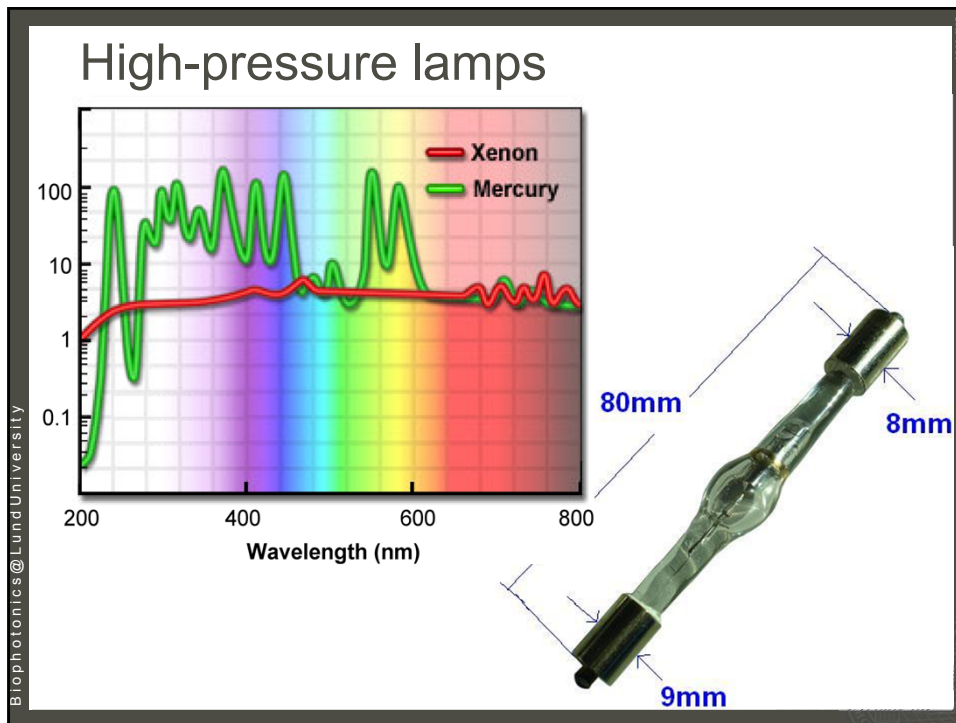
Blue shift Red shift

λ_0 λ

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Synchrotron Radiation

Electrons accelerated give rise to electromagnetic radiation – compare an antenna or *Bremsstrahlung* in nuclear physics



Storage ring with user beam lines

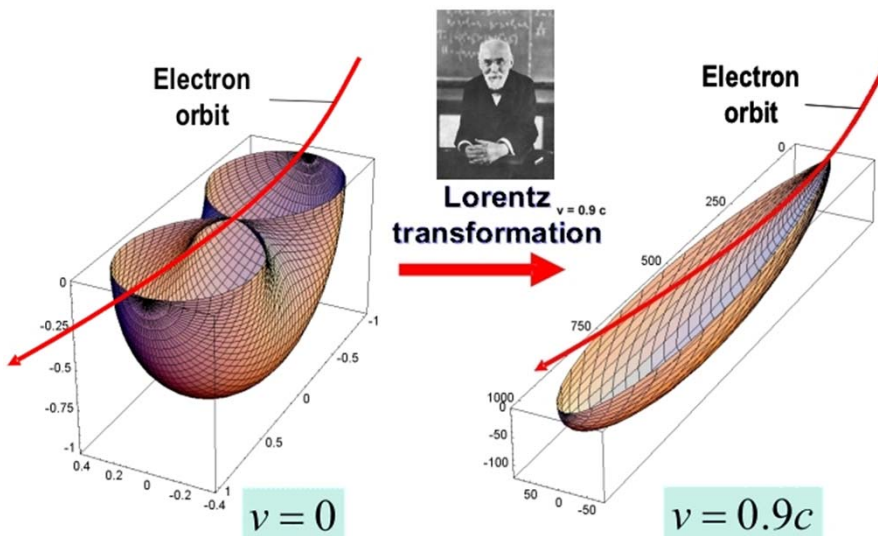
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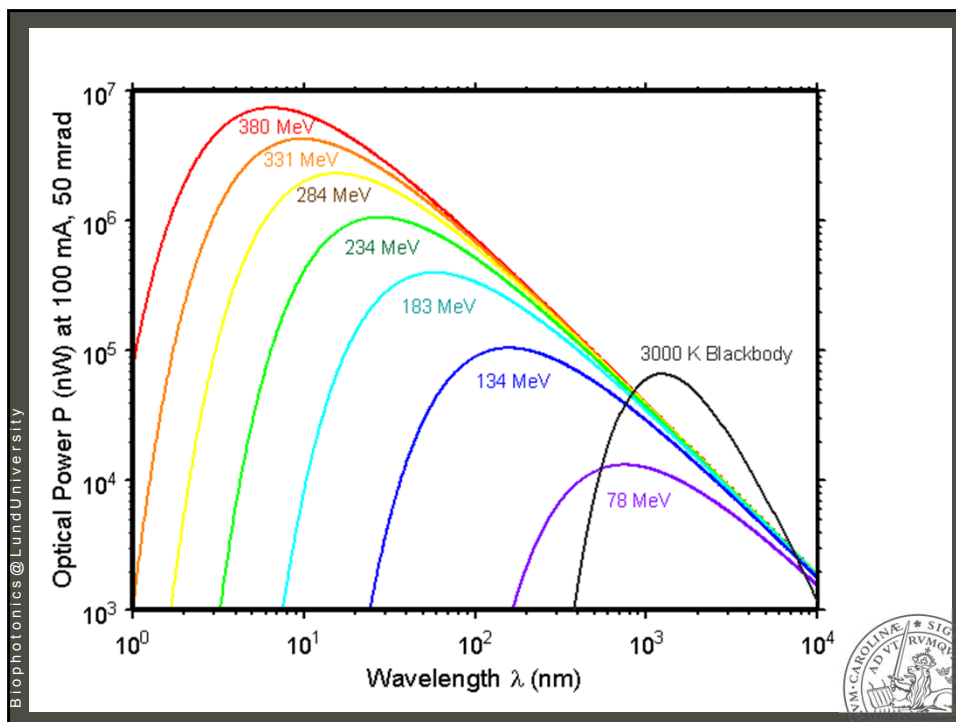
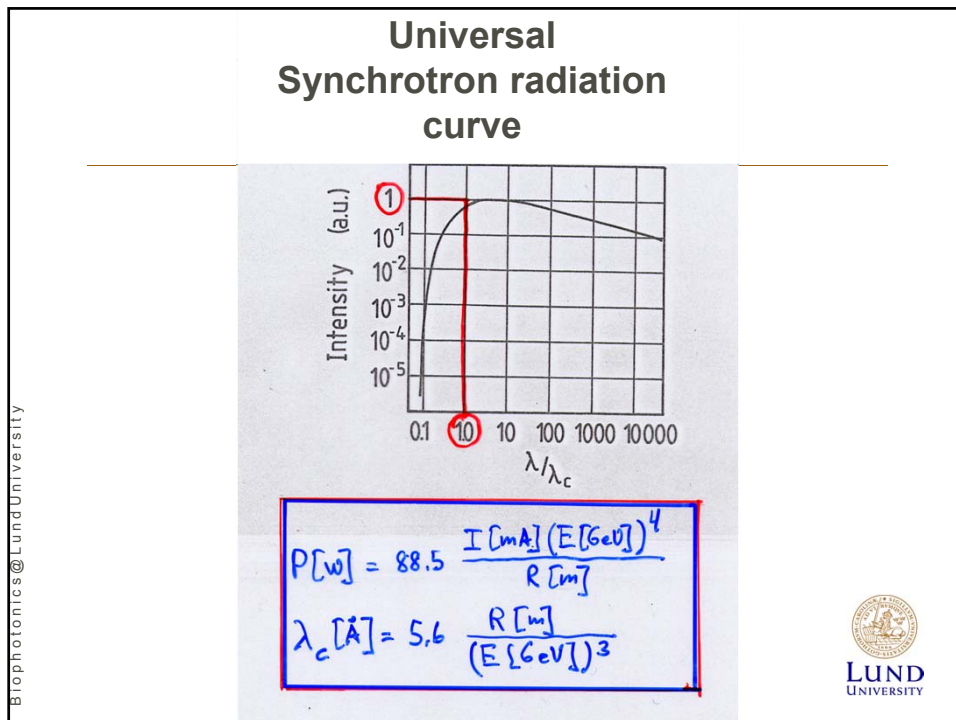
Synchrotron Radiation

For a dipole

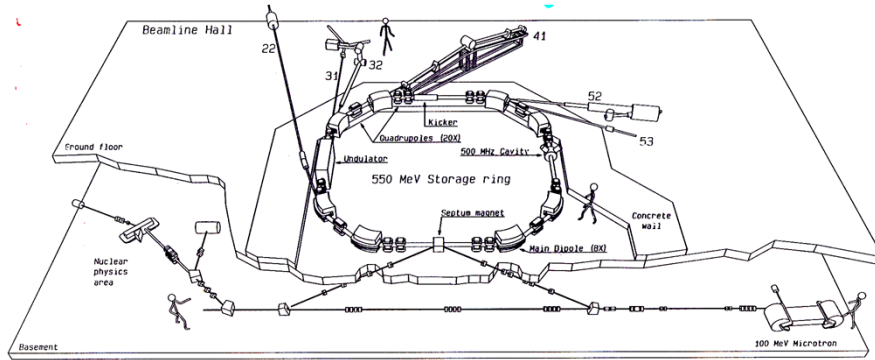
For relativistic electrons



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MAX-lab Synchrotron radiation storage ring, 550 MeV



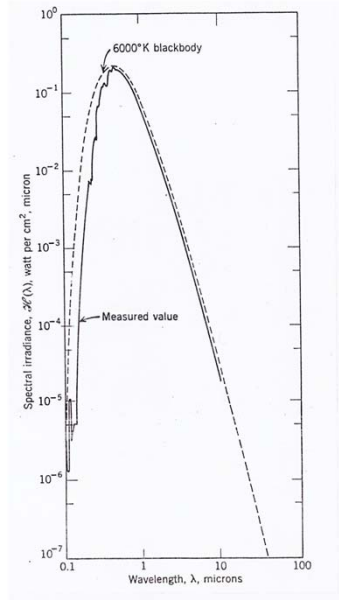
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Max IV, Lund - 3GeV Ring



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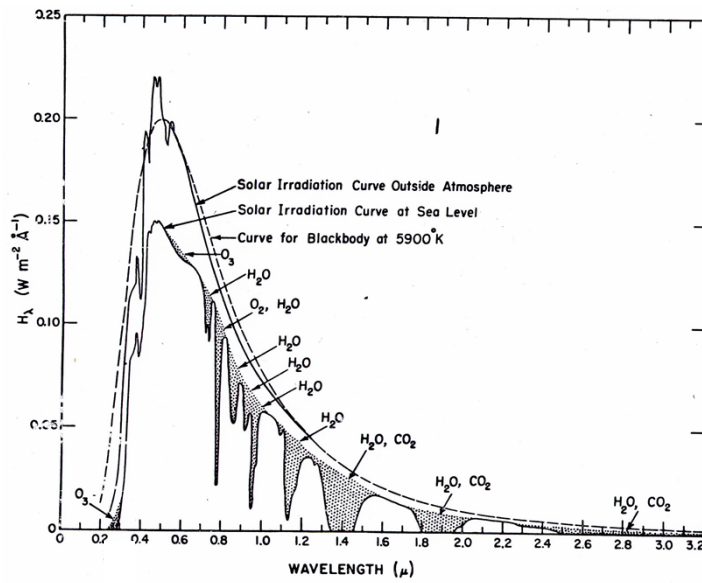
Spectral distribution of sunlight



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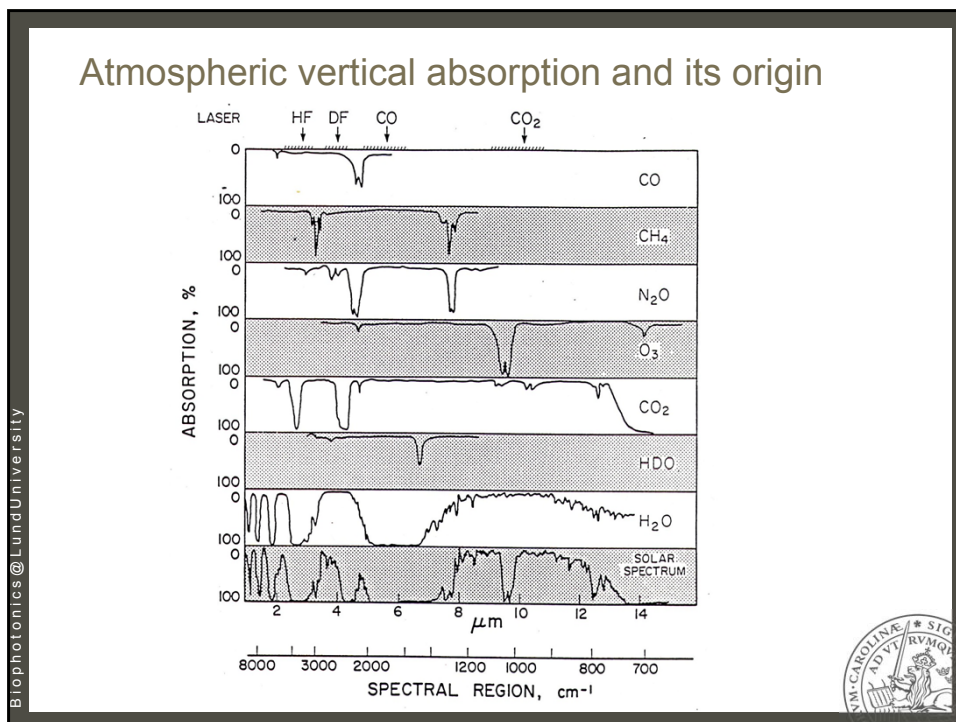
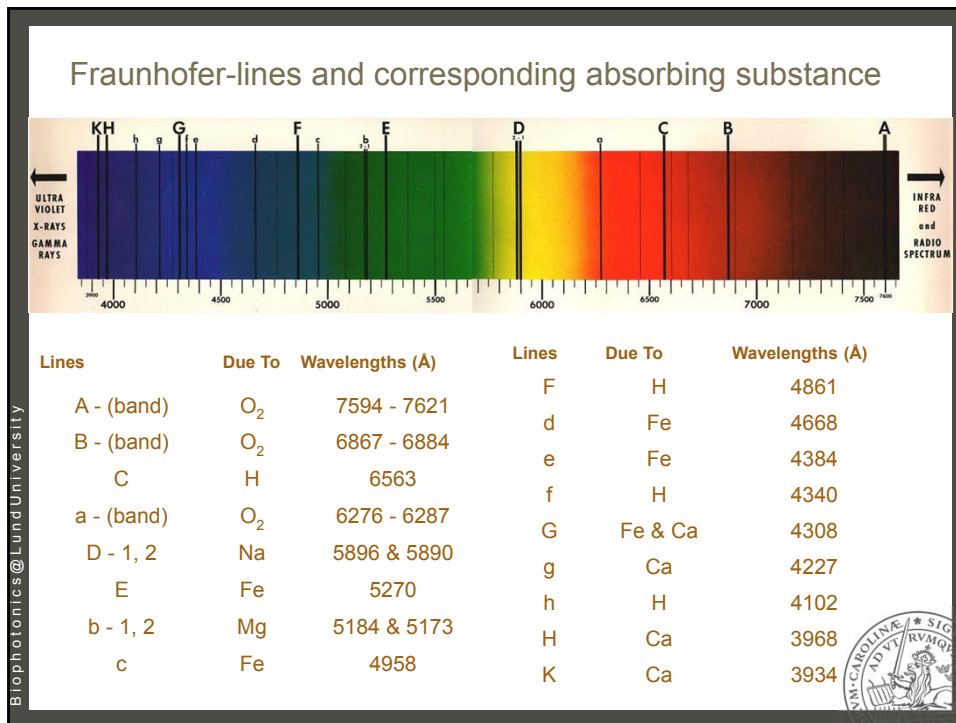


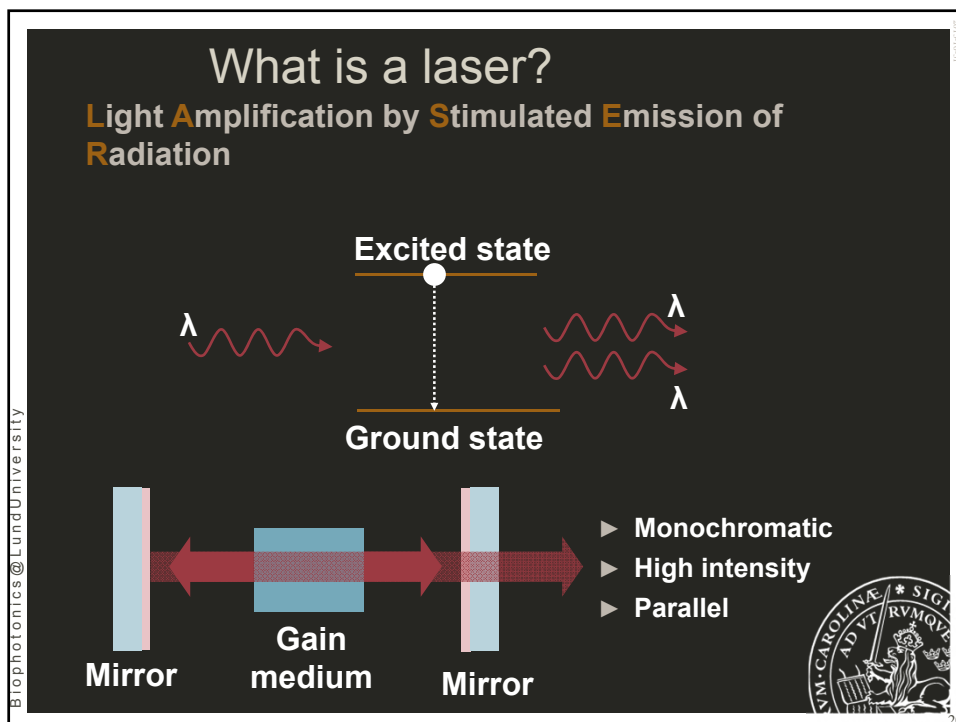
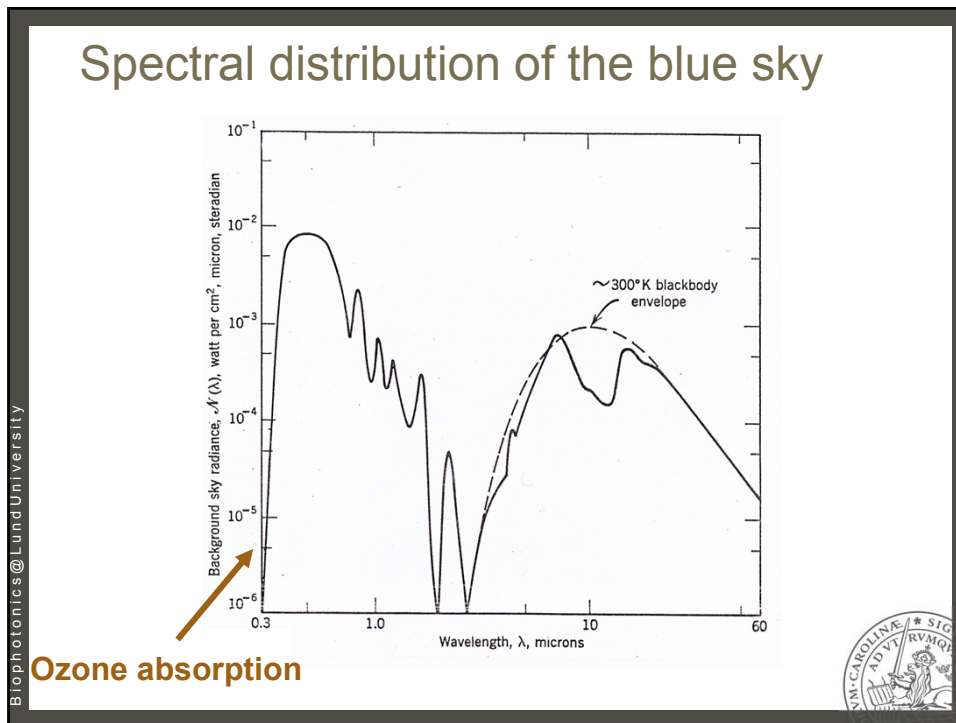
Sun spectrum



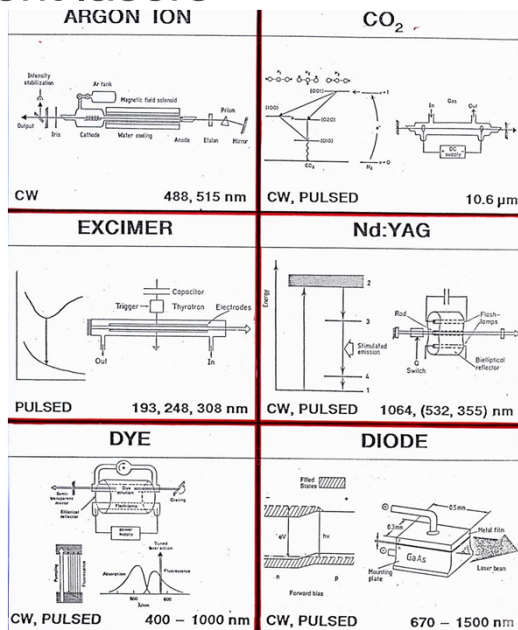
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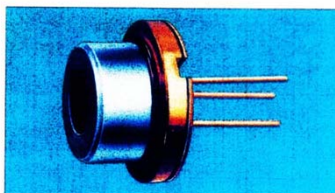
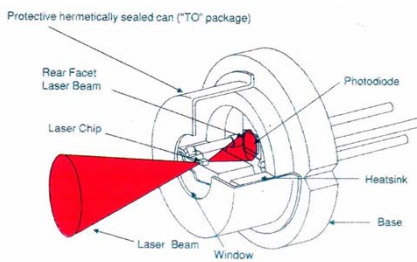
Different lasers



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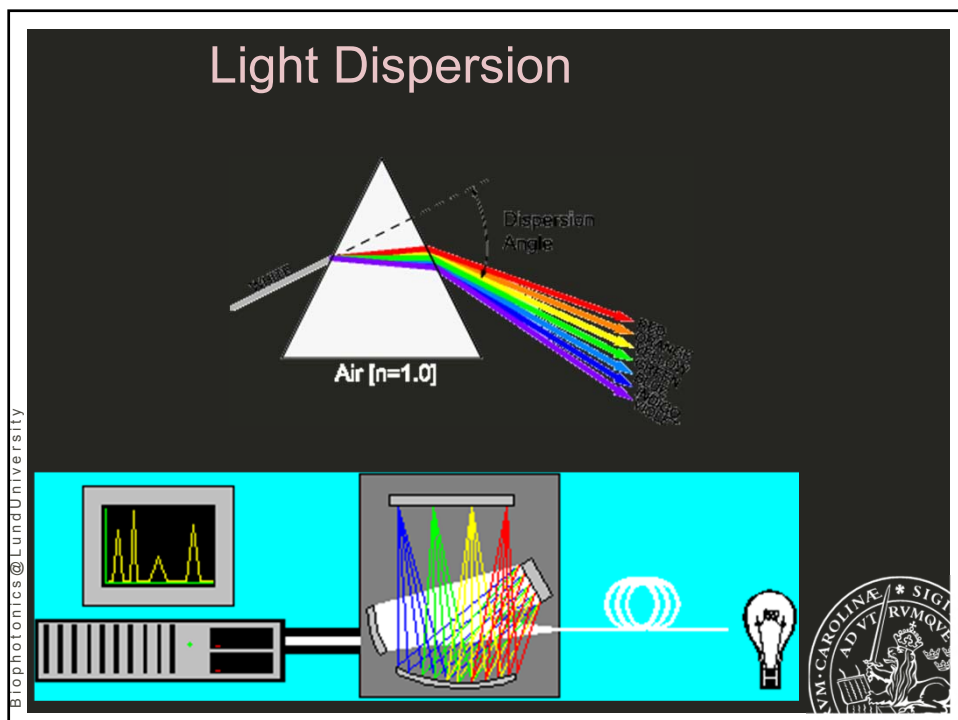
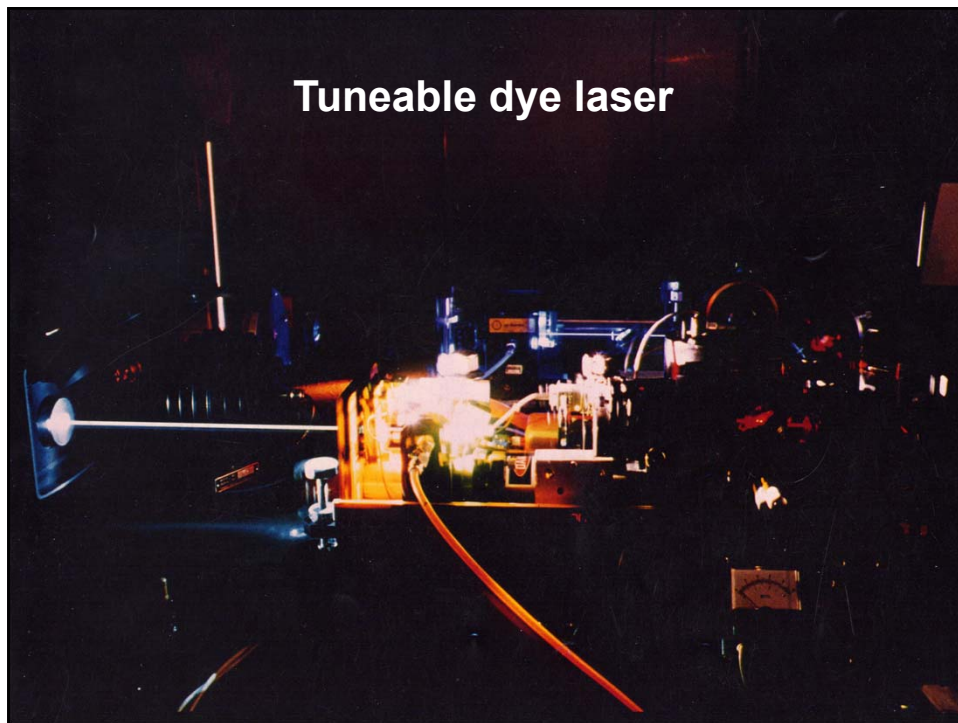
Laser Diodes



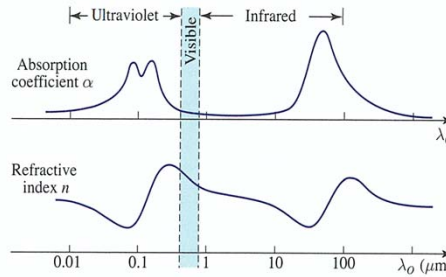
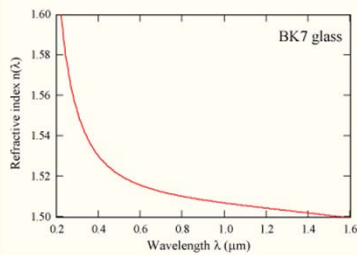
A typical compact disc style laser diode.

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The Prism Spectrometer



$$n^2 \approx 1 + \sum_i \chi_{0i} \frac{\nu_i^2}{\nu_i^2 - \nu^2} = 1 + \sum_i \chi_{0i} \frac{\lambda^2}{\lambda^2 - \lambda_i^2}$$

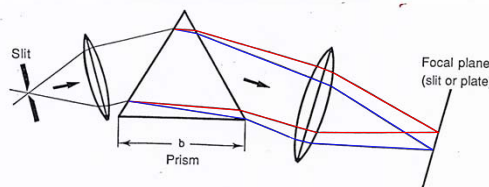
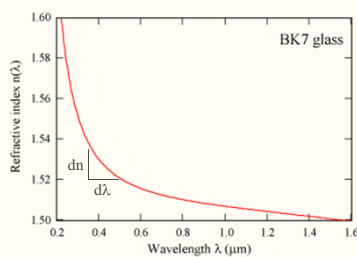
Fused silica

$$n^2 = 1 + \frac{0.6962\lambda^2}{\lambda^2 - (0.06840)^2} + \frac{0.4079\lambda^2}{\lambda^2 - (0.1162)^2} + \frac{0.8975\lambda^2}{\lambda^2 - (9.8962)^2}$$

W. Sellmeier, Annalen der Physik und Chemie 143, 271 (1871)

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The Prism Spectrometer

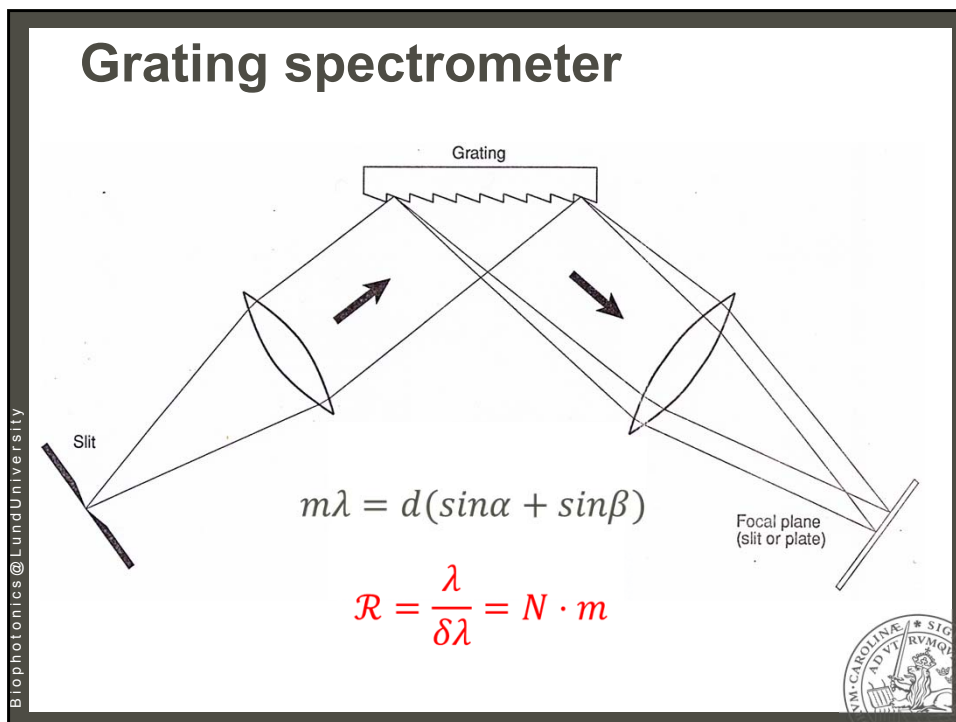
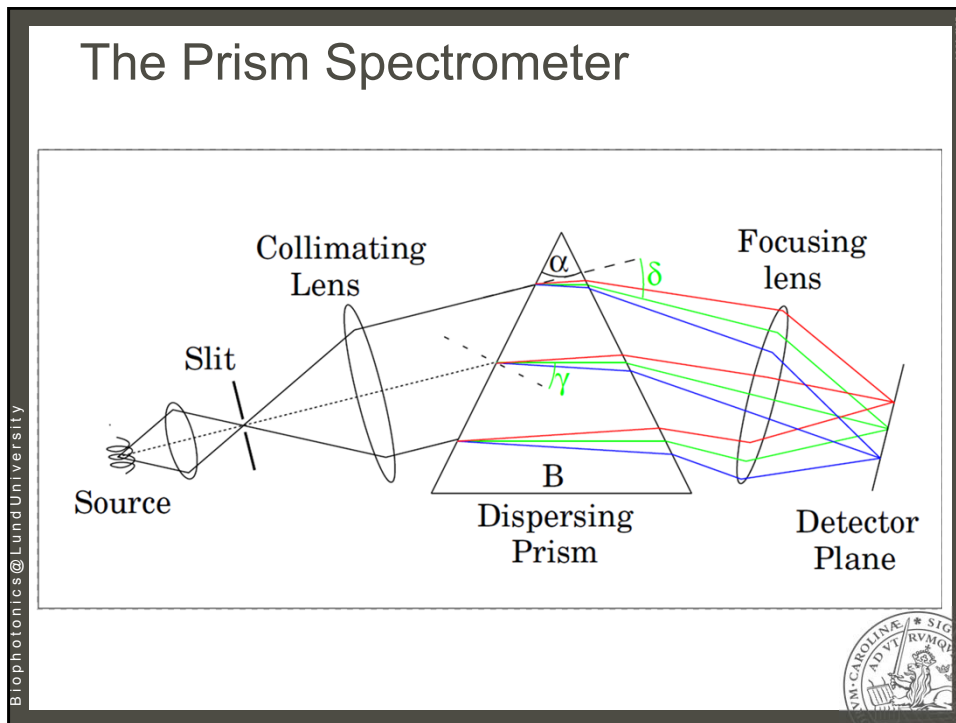


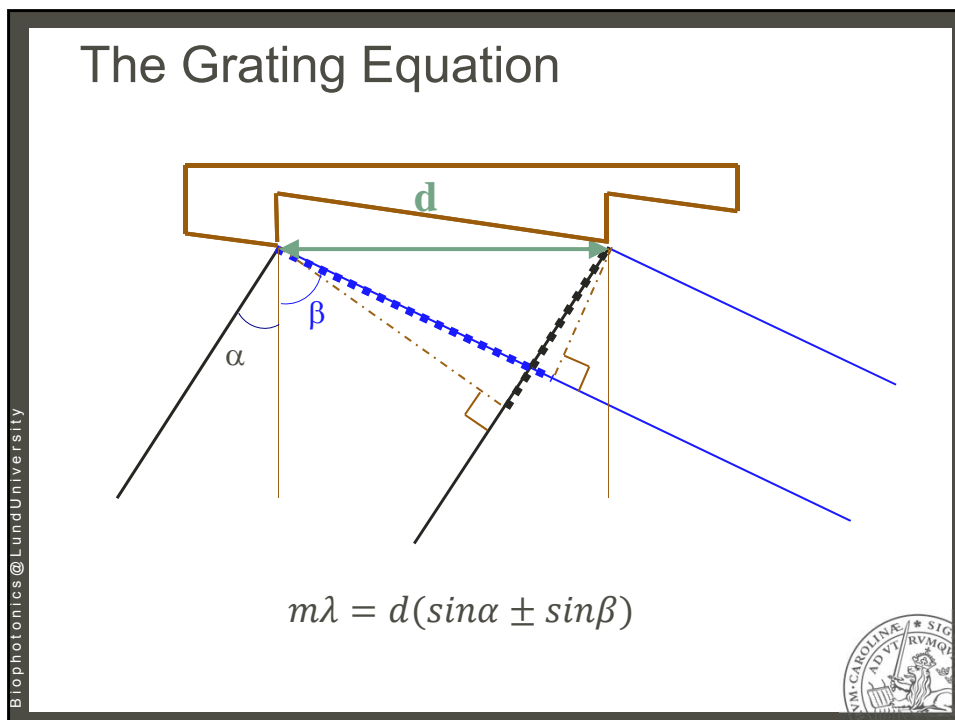
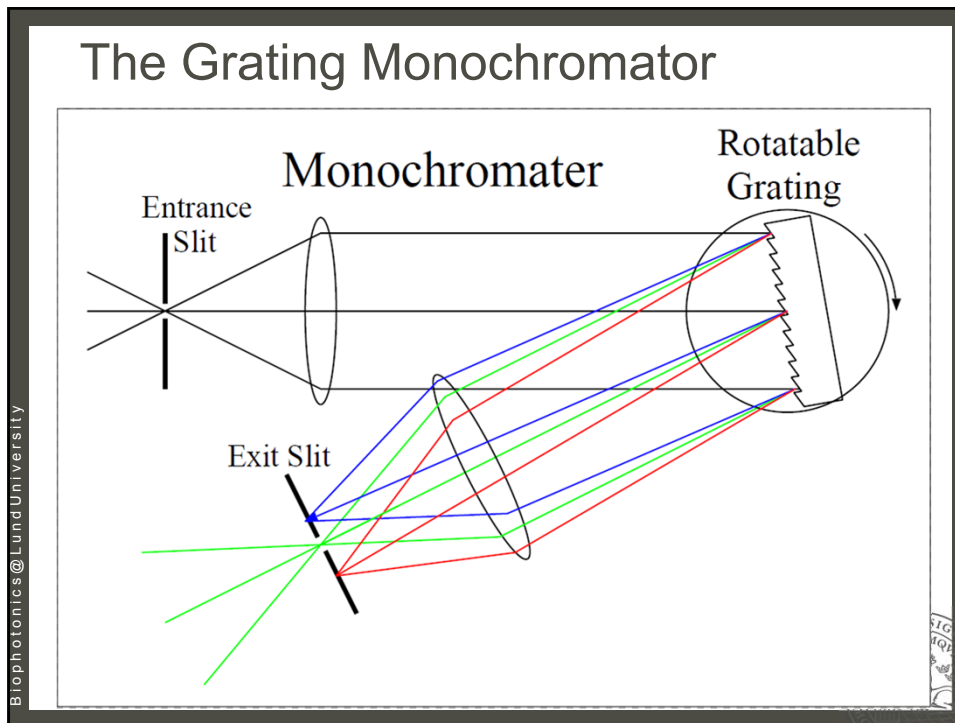
$$\mathcal{R} = \frac{\lambda}{\delta\lambda} = b \frac{dn}{d\lambda}$$

W. Sellmeier, Annalen der Physik und Chemie 143, 271 (1871)

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Fourier transform spectrometer

Michelson interferometer

I Interferogram

B Spectrum

$I(\Delta) = I_0 \cos^2 \frac{\phi}{2}$
 $\phi = \frac{\Delta}{\lambda} 2\pi = \frac{\Delta}{c} 2\pi \nu$ for single frequency. Superposition princ. →

$I(\Delta) = \int_0^\infty B(\nu) \cos^2 \left(\frac{\Delta}{c} \pi \nu \right) d\nu = \frac{1}{2} \int_0^\infty B(\nu) [1 + \cos \left(\frac{\Delta}{c} 2\pi \nu \right)] d\nu$

$f(\Delta) = \frac{1}{2} \int_0^\infty B(\nu) \cos \left(\frac{\Delta}{c} 2\pi \nu \right) d\nu$ Interferogram

$B(\nu) \propto \int_{-\infty}^{\infty} f(\Delta) \cos \left(\frac{\Delta}{c} 2\pi \nu \right) d\Delta$ Spectrum

APODIZATION Folgt dem Zerfall } **advantage**

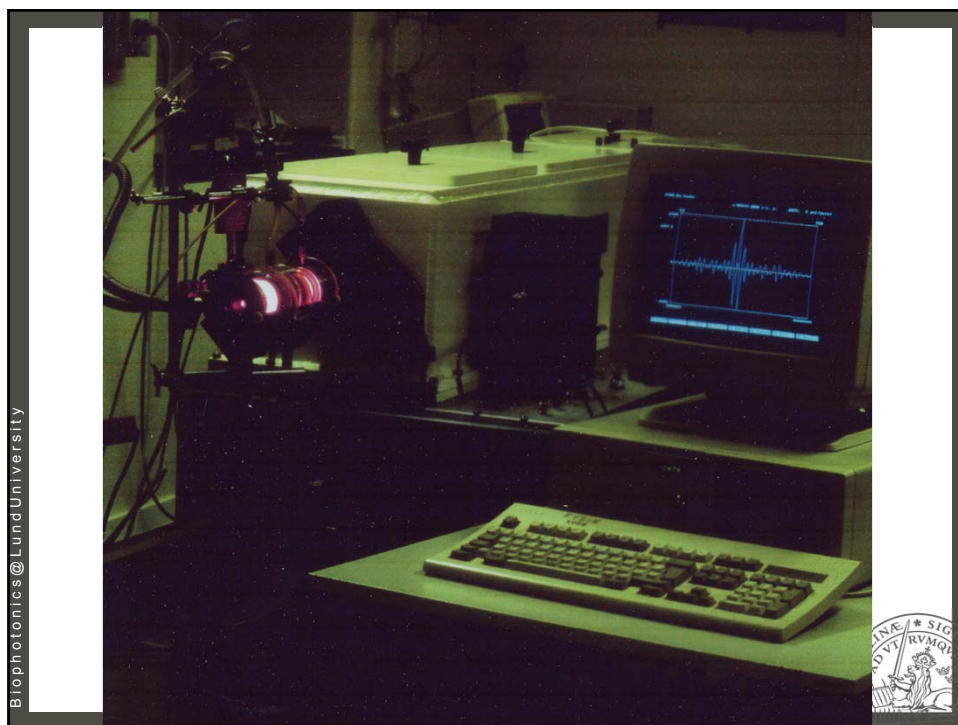
6-32

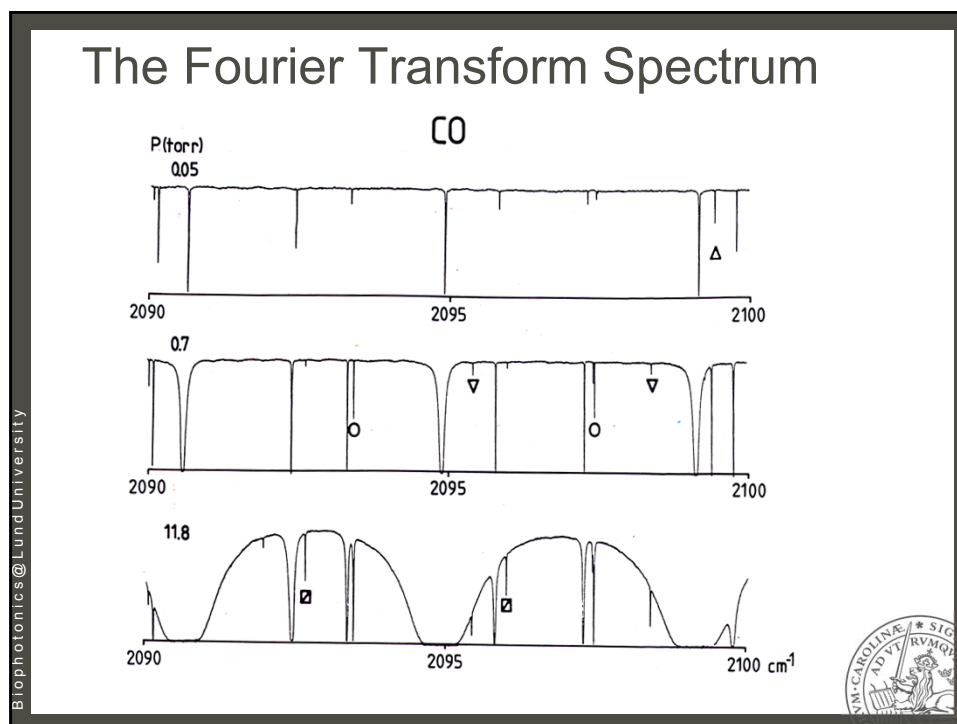
FTIR

Detector Output

Mirror Retardation

Spectra





Photographic films

Emulsion deposited on plastic or glass
 microscopic crystals of *silver halide* in gelatin AgCl;
 AgBr and AgI

When exposed to light, latent image forms
 silver ions reduced to metallic silver

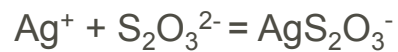
Developer is a reducing agent e.g. quinol
 reduces ions to metal
 preferentially where granule contains nucleus of
 metallic silver

$$2\text{Ag}^+ + \text{C}_6\text{H}_4(\text{OH})_2$$


Film

Dark parts are where light has fallen
Negative

Fixing film dissolves silver halide



Positive print

shine light through negative onto paper coated
with silver halide.

Developed & fixed

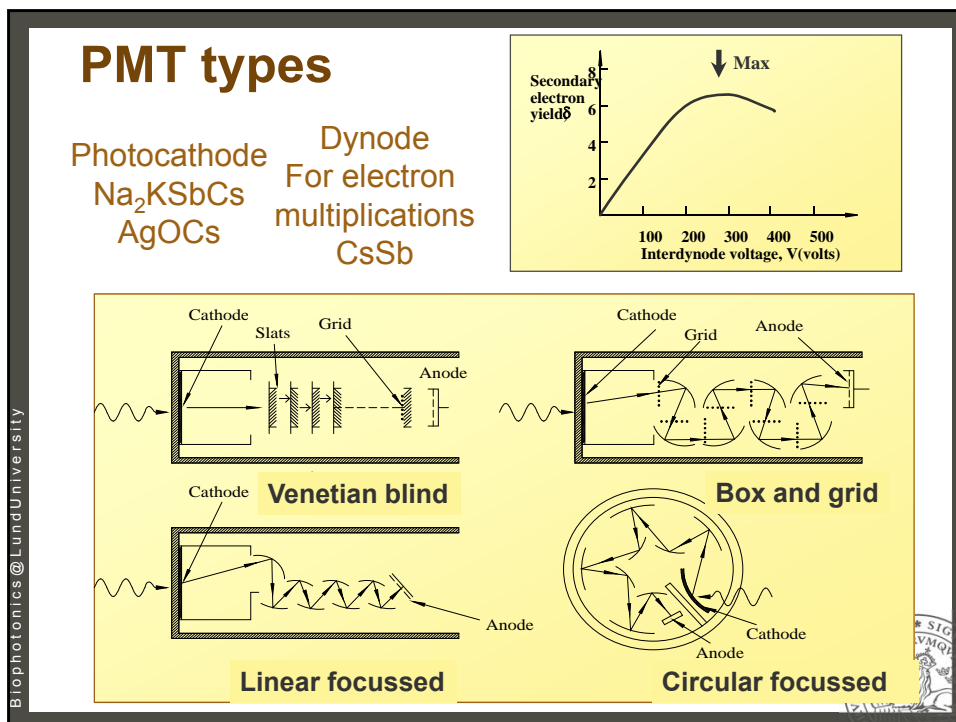
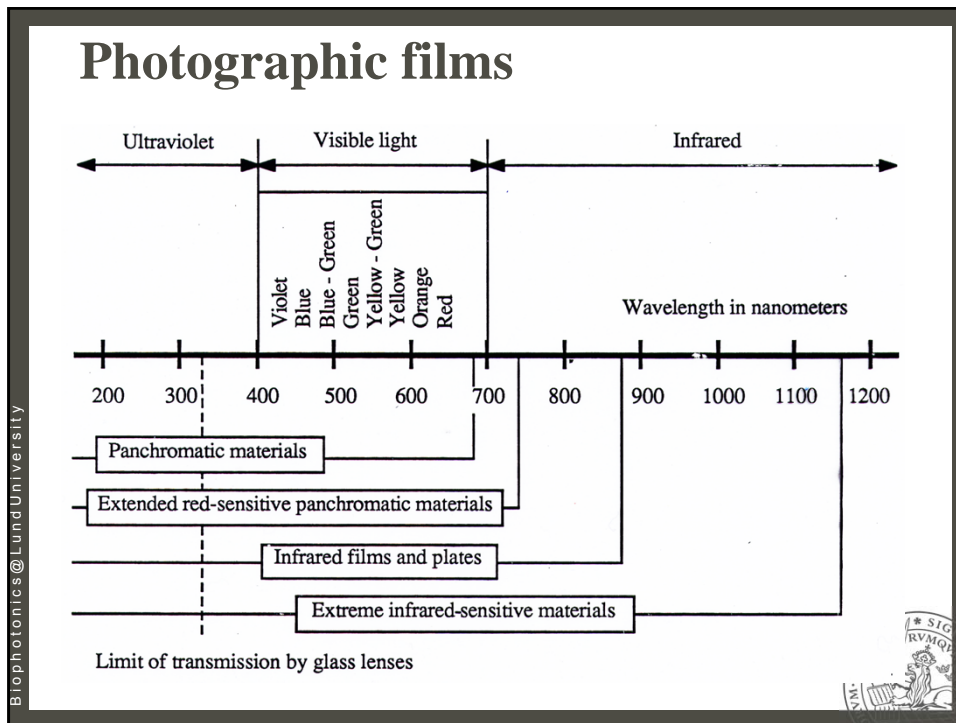


Colour film



Colour films use 3 layers of silver halide
layers sensitised with different colour dye
layers separated by colour filters





CCD Cameras

Light focused onto CCD chip

Each capacitor results in one *pixel*

a pixel is a picture element with no internal detail

Number of pixels on the chip determines resolution of camera

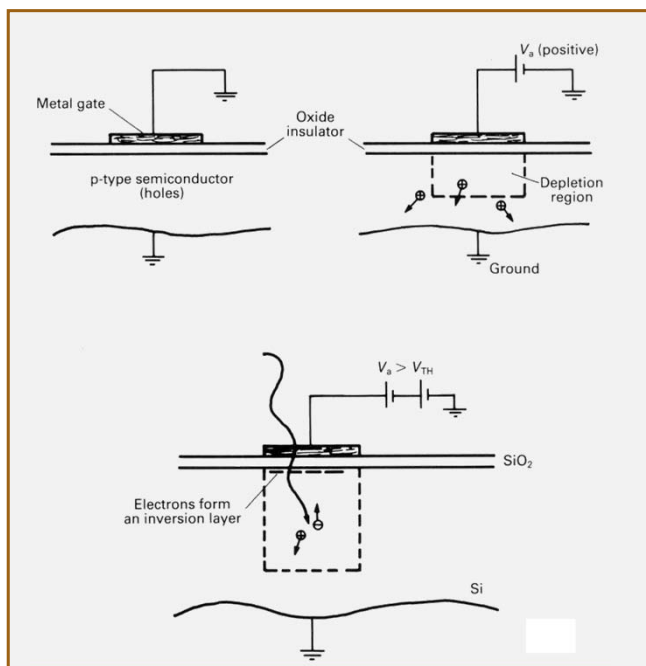
Commonly 3-5 Megapixels

up to 13 Mpixels

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CCD

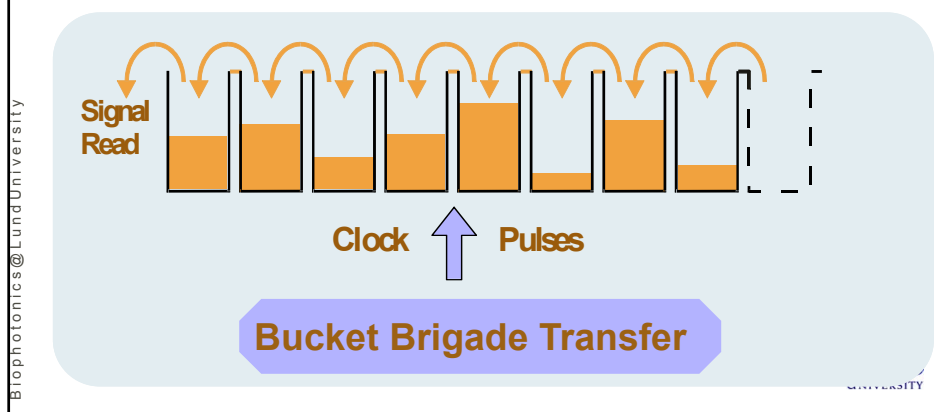


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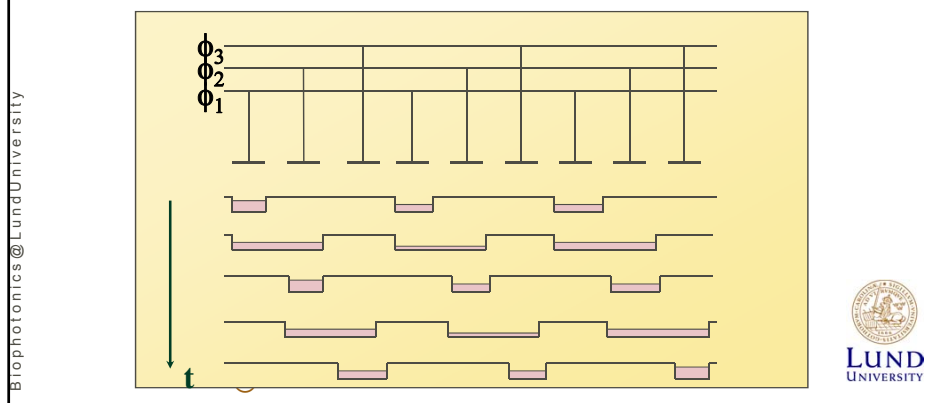
Readout of a CCD

Charge moved through chip by *bucket brigade transfer*
charge on final capacitor is read off

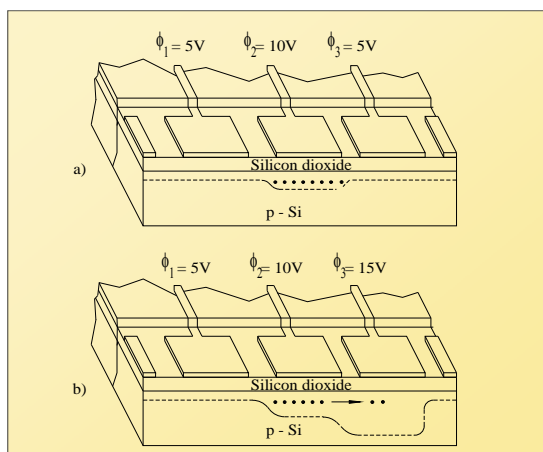


CCD camera readout

Schematic potential well diagram for transport of photo-charges
along one direction in the CCD-chip



CCD - Charge Coupled Device



The information is read out by stepwise transferring by manipulating the potential wells.



CMOS Cameras

Recently, CMOS's have become more used in cheap cameras instead of CCD's

CMOS stands for Complementary Metal Oxide Semiconductor.

Each pixel has a circuit on the chip to convert the charge to a voltage. The camera often also include an analogue-to-digital converter.



Applications of CMOS cameras

Advantages

- Compact
- Cheap
- Low power
- Fast readout
- Plug & Play

Disadvantages

- Nonuniform
- Dark areas

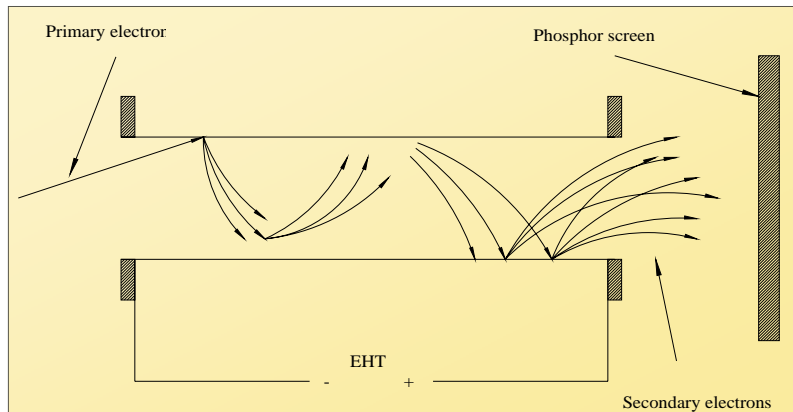


Infrared array detectors



InGaAs arrays are required for spectral analysis in the near to middle infrared (0.9 - 2.6 μm). High performance multi-channel detectors for this spectral region are a **recent development**. The vast majority of detectors sold today are based on silicon technology which only operates on the shorter wavelength side of the NIR region (< 1.1 μm).

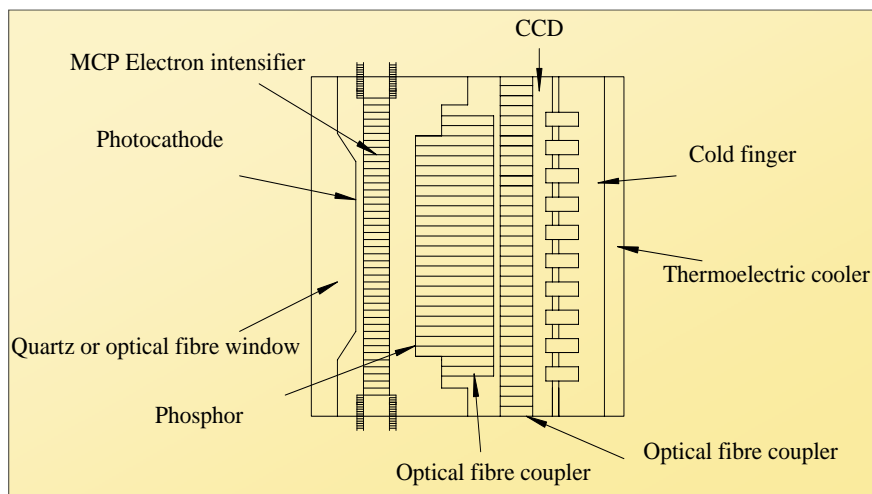
Micro-channel plate image intensifier



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MCP



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Image Intensified CCD Cameras

The features of the new ICCD range include:



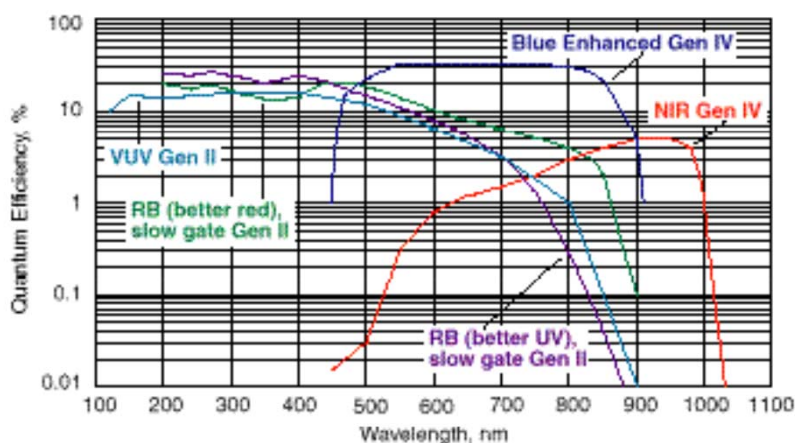
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- Single photon sensitivity
- 15 bit dynamic range in a single scan
- Gating widths in the order of ns (ps with special image intensifiers)
- Compact head and controller
- A wide selection of photocathode options
- Easy interfacing to delay generators and lasers
- Modern sealed design - no nitrogen flush
- Thermoelectric air-cooled convenience



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Spectral response of ICCDs



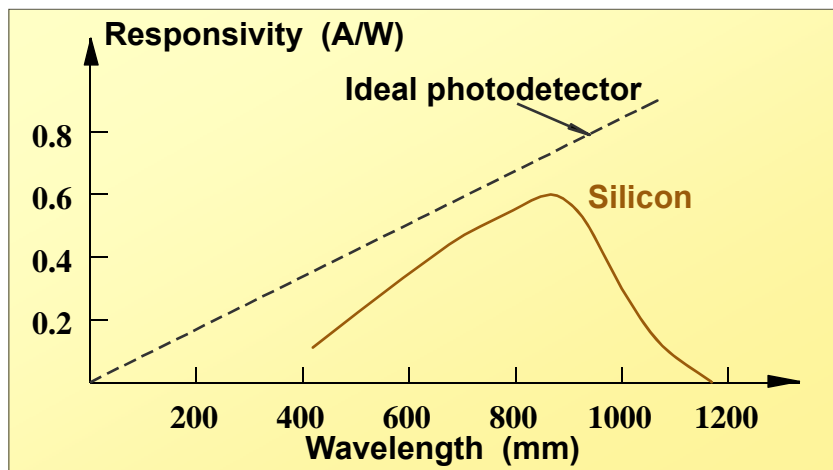
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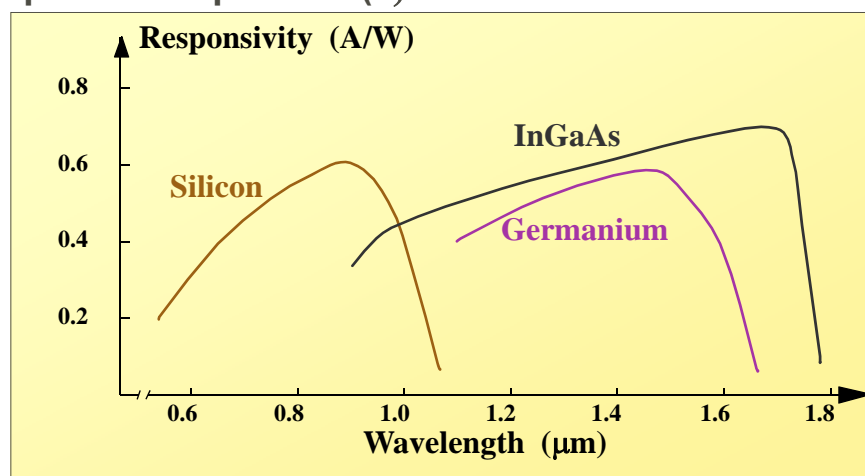
Choosing a detector - detector parameters

► Spectral response $R(\lambda)$



Choosing a detector - detector parameters

► Spectral response $R(\lambda)$



Choosing a detector - detector parameters

► Spectral response $R(\lambda)$

- Sensitivity $S(\lambda) = \text{signal/optical power [V/W] or [A/W]}$
- Signal-to-noise = $S/N = V_{\text{signal}}/V_{\text{noise}}$

NEP = Noise Equivalent Power ==> if $P = \text{NEP}$ then $S/N = 1$

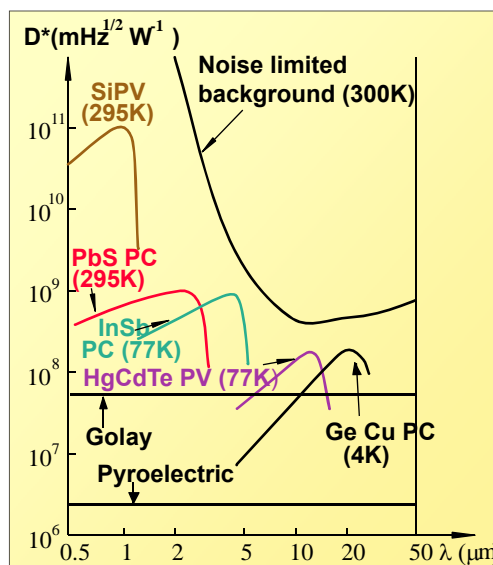
$$\text{NEP}^* = \text{NEP} / \sqrt{A \cdot \Delta f} \quad \text{Normalized NEP}$$

$$D^* = 1 / \text{NEP}^* \quad \text{Specific detectivity [m}\sqrt{\text{Hz}} / \text{W]}$$

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Example



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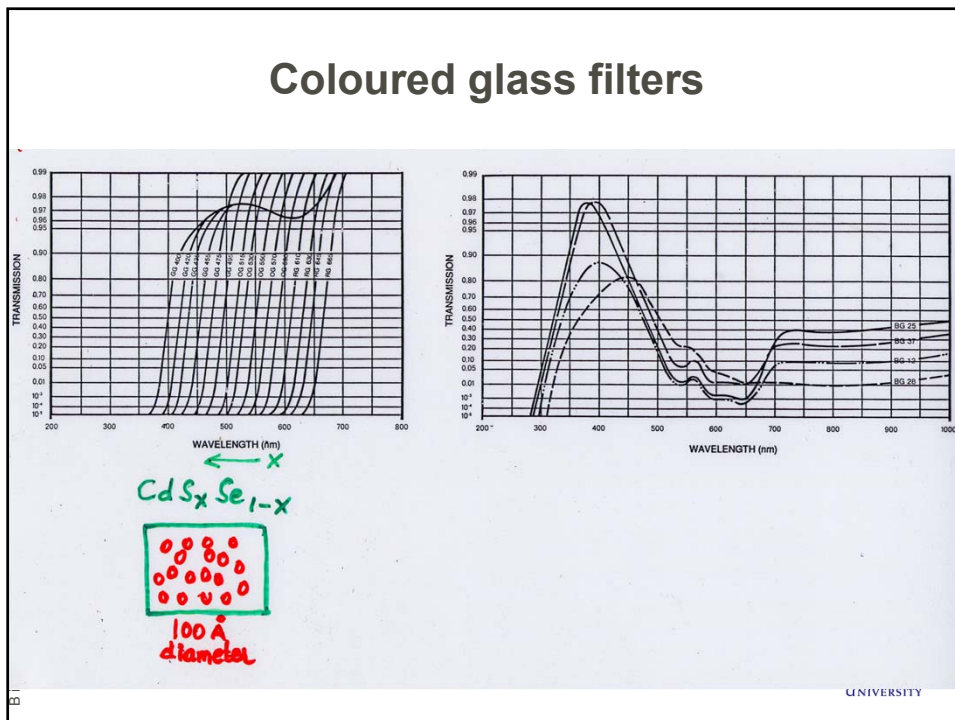
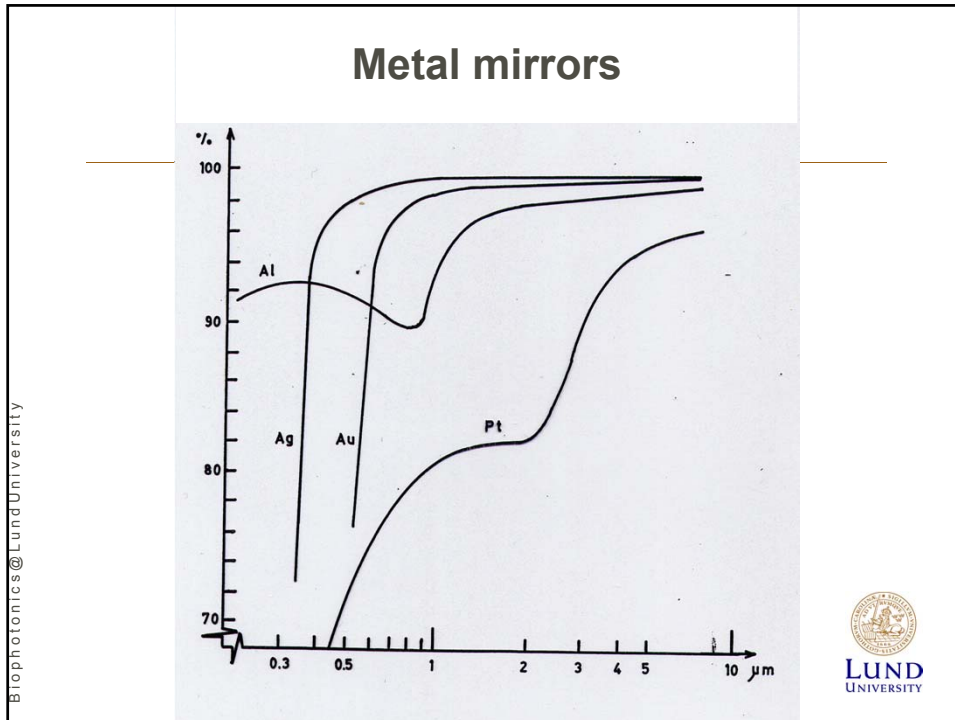
Calculate the minimum detectable signal at $\lambda = 4$ mm. Choose an InSb detector. Size 100 mm^2 , bandwidth 1 kHz

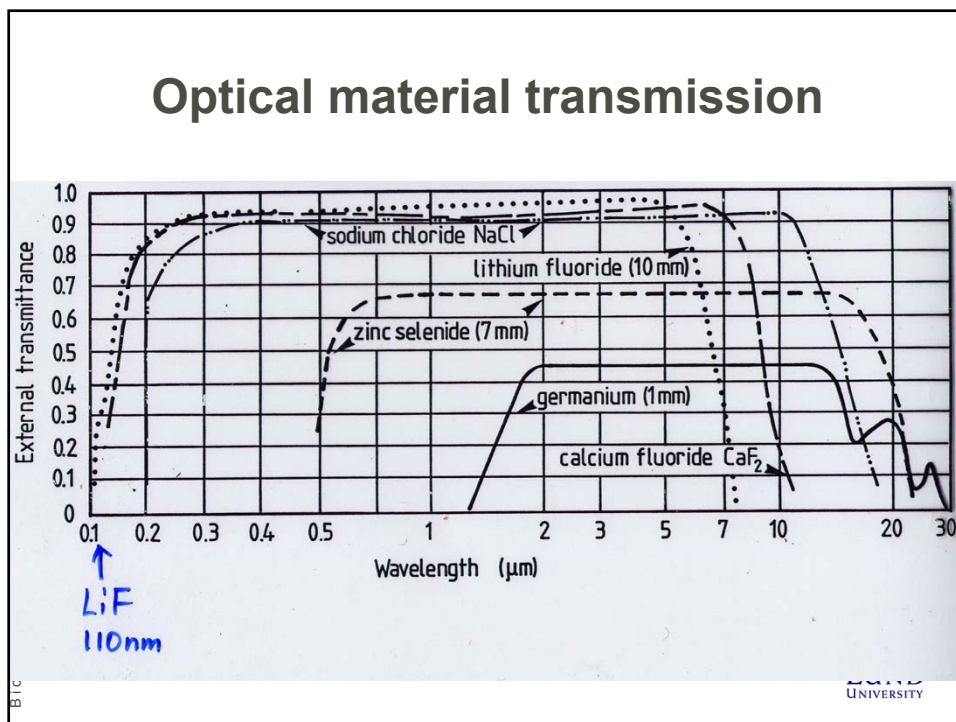
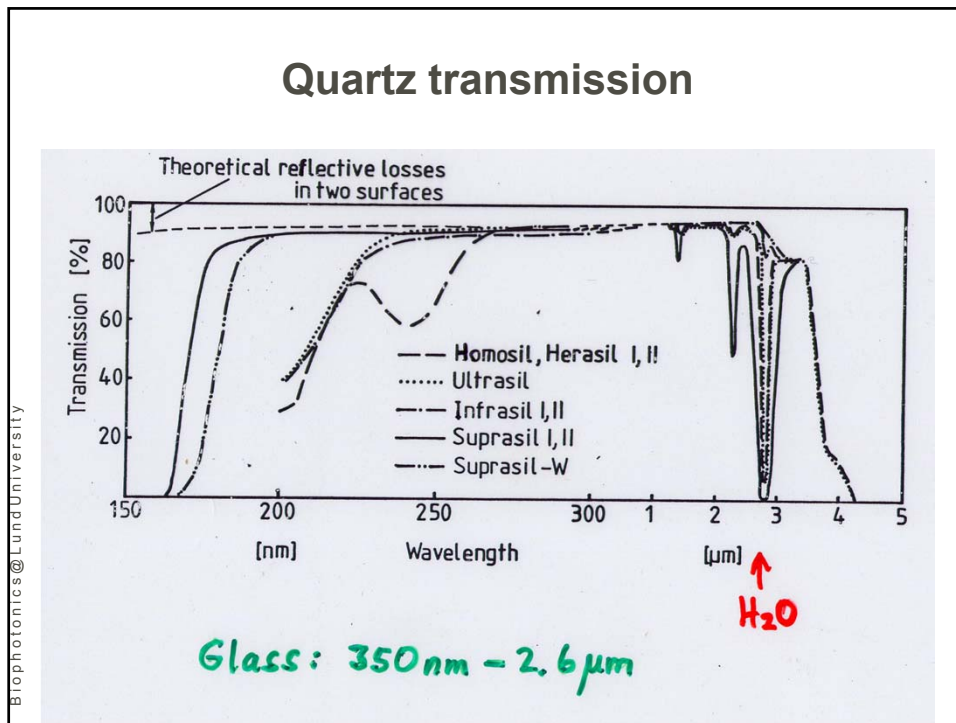
$$D^* = \frac{1}{\text{NEP}^*} = \frac{\sqrt{A \Delta f}}{\text{NEP}} \Rightarrow$$

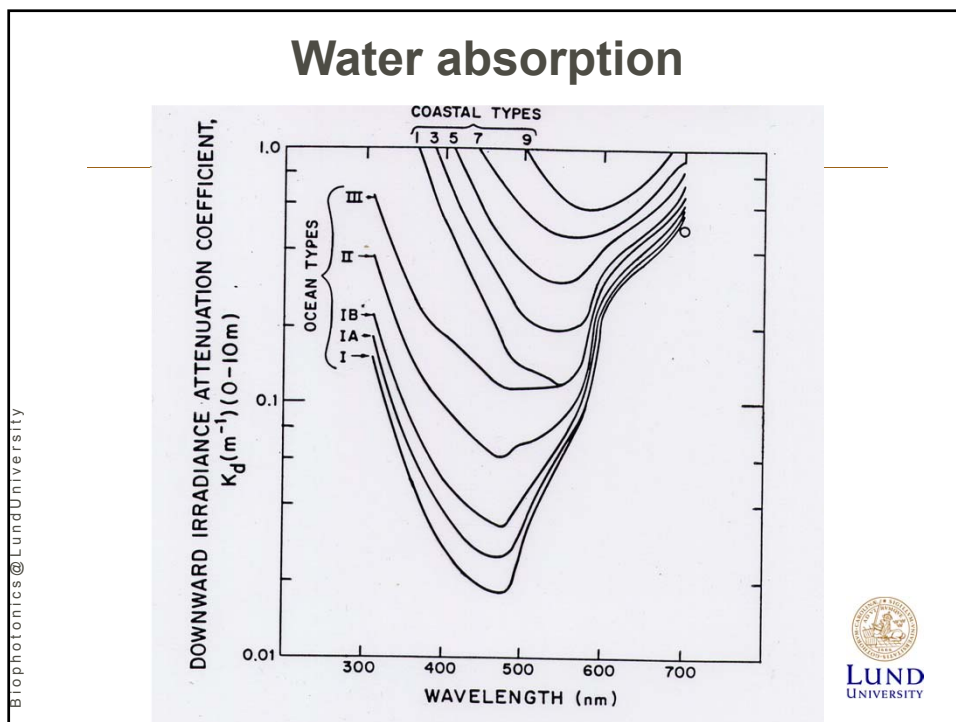
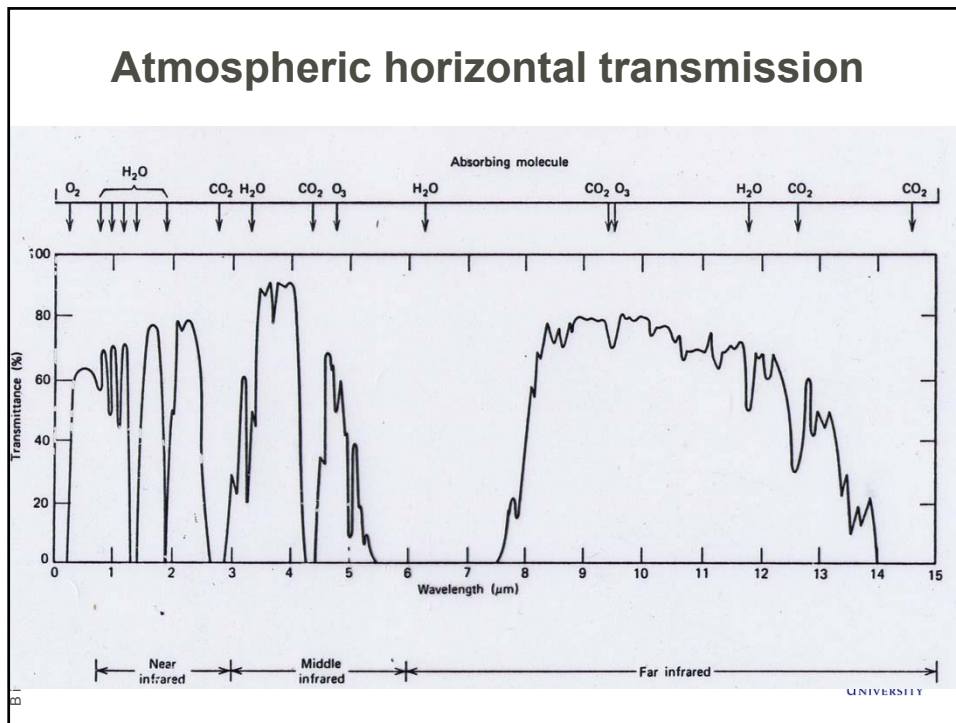
$$\text{NEP} = \frac{\sqrt{A \Delta f}}{D^*} =$$

$$\frac{\sqrt{100 \cdot 10^{-6} \cdot 1000}}{10^9} = 3 \cdot 10^{-10} \text{ W}$$









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The Beer-Lambert law

$$\ln \frac{P_0}{P_t} = k_1 \cdot b \cdot c$$

$$A = \log_{10} \frac{P_0}{P_t} = 0.434 \ln \frac{P_0}{P_t}$$

$A \propto c$

$$T = \frac{P_t}{P_0}$$

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Atomic absorption spectroscopy

OK!

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